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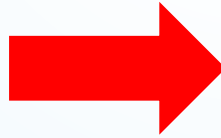
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A Companion to Science, Technology, and Medicine in Ancient Greece and Rome

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Summary

The history of ancient Greek and Roman **astronomy** can traditionally be divided into three broad periods, Presocratic speculation, the homocentric spheres model of Eudoxus and the epicyclic model of Ptolemy. This in itself makes a decision on how to approach Greek and Roman **astronomy**, prioritizing the development of models of the heavens, Alternatively, one might focus on observation, practices or the practical application of **astronomy**. Here we concentrate on models but not to the exclusion of other approaches. We also challenge the idea that Presocratic **astronomy** was largely speculative. To treat **astronomy** on its own is also to some extent to impose a modern disciplinary boundary on ancient work which did not recognize a sharp boundary with cosmology, or in some cases with astrology.

Chapter

Data-Intensive Methods in Astronomy

Thomas D. Kitching, Robert G. Mann, Laura E. Valkonen, Mark S. Holliman, Alastair Hume, Keith T. Noddle

The DATA Bonanza: Improving Knowledge Discovery in Science, Engineering, and Business

First published: 09 April 2013

Summary

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This chapter first briefly outlines the concept of the virtual observatory as a science response to the growing wealth of **astronomical** data and the increasing requirement to answer scientific questions using **astronomical** data from multiple sources. Then, it describes two examples of data-intensive applications from **astronomy**. The first example centers on running photometric classification algorithms on two existing sky survey databases to generate a catalog of quasars, while the second example outlines the derivation of constraints on cosmological parameters via weak gravitational lensing, which is one of the main science drivers for the next generation of sky surveys. This pair of examples illustrates some of the problems faced by modern **astronomy**, which require the development of new data-intensive methods and the availability of a computational infrastructure on which to run them.

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The Dawn of X-Ray Astronomy (Nobel Lecture)

Riccardo Giacconi

ChemPhysChem | Volume 4, Issue 7

First published: 07 July 2003

Abstract 

Abstract

Riccardo Giacconi joined the American Science and Engineering Corporation (AS & E) after leaving Princeton University in 1959, and in 1962 his group there detected the first extrasolar X-ray source. Prof. Giacconi was subsequently responsible for the launch and use of the satellite UHURU (1970) and the EINSTEIN observatory (1978). He was appointed Associate Director of the High Energy Astrophysics Division of the Harvard-Smithsonian Center for Astrophysics in 1973 and was also appointed Professor of Astronomy at Harvard University that same year. In 1981 he became the first Director of the Space Telescope Science Institute and was also appointed Professor in the Department of Physics and Astronomy at Johns Hopkins University. In 1992 he was appointed Director General of the European Southern Observatory, an intergovernmental organization of eight nations. Prof. Giacconi is currently President of Associated Universities, Inc., and Research Professor at Johns Hopkins University. He was awarded the Wolf Prize in 1987, and the Nobel Prize for Physics in 2002.

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The Dawn of X-Ray Astronomy (Nobel Lecture)[†]

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1.0 Introduction

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Active Galactic Nuclei

Author(s): Dr. Volker Beckmann, Dr. Chris Shrader
First published: 29 August 2012
Print ISBN: 9783527410910 | Online ISBN: 9783527666829 | DOI: 10.1002/9783527666829
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About this book

This AGN textbook includes phenomena based on new results in the X-Ray domain from new telescopes such as Chandra and XMM Newton not mentioned in any other book. Furthermore, it considers also the Fermi Gamma Ray Space Telescope with its revolutionary advances of unprecedented sensitivity, field of view and all-sky monitoring. Those and other new developments as well as simulations of AGN merging events and formations, enabled through latest super-computing capabilities.

The book gives an overview on the current knowledge of the Active Galactic Nuclei phenomenon. The spectral energy distribution will be discussed, pointing out what can be observed in different wavebands and with different physical models. Furthermore, the authors discuss the AGN with respect to its environment, host galaxy, feedback in galaxy clusters, etc. and finally the cosmological evolution of the AGN phenomenon.

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